

Coming attraction: Brodie, Friday, 16:10

How to cool a neutron star *very* fast



Liam Brodie & RDP, [2501.02055](#) , [PRL 135](#) .

SU(2) parity doubled model:

D. Zschiesche, L. Tolos, J. Schaffner-Bielich & RDP, [0608044](#)

L. Brodie, R. Negreiros, J. Steinheimer, V. Dexheimer, & RDP, [2603.06789](#)

Punchline: in parity doubled models, for *limited* region of density, in *heavy* stars,
Urca with parity doubled nucleons opens up, *neutrino emission increases by 10^3 !*

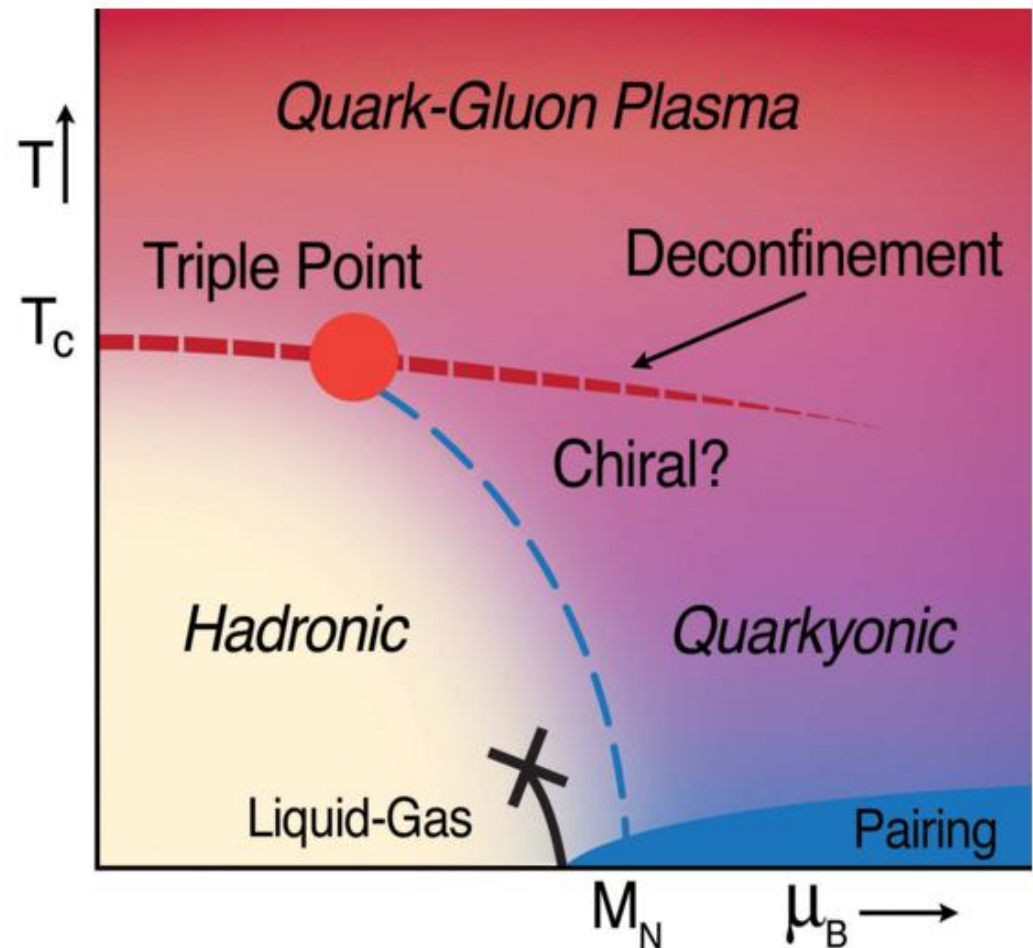
“Quarkyonic” phase diagram

When $\# \text{ colors} \gg \# \text{ flavors}$, gluons (obviously) dominate, “*quarkyonic*”

L. McLerran & RDP, [0706.2191](#)

At “moderate” μ , chiral symmetry restoration happens *before* deconfinement
Need model of chirally symmetric baryons.

Andronic...McLerran, RDP +...
[0911.4806](#)



Chirally symmetric, *massive* baryons

Massless quarks have chiral symmetry. For *massive* baryons? Detar and Kunihiro '89:

Usual nucleons, $N(939)$, $J^P = \frac{1}{2}^+$, *and* their parity partners, $N^*(1535)$, $J^P = \frac{1}{2}^-$

$$\mathcal{L}_N = \bar{\psi}_1 \not{\partial} \psi_1 = \bar{\psi}_{1,L} \not{\partial} \psi_{1,L} + \bar{\psi}_{1,R} \not{\partial} \psi_{1,R} \quad \psi_{L,R} = \frac{1 \pm \gamma_5}{2} \psi$$

$$\mathcal{L}_{N^*} = \bar{\psi}_2 \not{\partial} \psi_2 = \bar{\psi}_{2,L} \not{\partial} \psi_{2,L} + \bar{\psi}_{2,R} \not{\partial} \psi_{2,R}$$

Duh. Big deal. But because N^* 's parity differs, so do chiral transformations:

$$\psi_{1;L,R} \rightarrow U_{L,R} \psi_{1;L,R} ; \quad \psi_{2;L,R} \rightarrow U_{R,L} \psi_{2;L,R}$$

Then can form chirally *symmetric* mass term for the N 's *and* N^* 's:

$$\mathcal{L}_m = m_0(\bar{\psi}_{1,L} \gamma_5 \psi_{2,R} + \bar{\psi}_{1,R} \gamma_5 \psi_{2,L} + L \leftrightarrow R) = m_0(\bar{\psi}_1 \gamma_5 \psi_2 - \bar{\psi}_2 \gamma_5 \psi_1)$$

Parity Doubled baryons

Couple these nucleons to $SU(2)_L \times SU(2)_R = O(4)$ invariant fields $\phi = (\sigma, \pi)$, along with vector mesons, ω_μ and ρ_μ :

$$\begin{aligned}\mathcal{L}_{PD'd} = & \bar{\psi}_1 (\not{\partial} - 2g_1 (\sigma + i\gamma_5 \vec{\pi} \cdot \vec{\tau}) - g_\omega \not{\omega} - g_\rho \not{\rho} \cdot \vec{\tau}) \psi_1 \\ & + \bar{\psi}_2 (\not{\partial} - 2g_2 (\sigma - i\gamma_5 \vec{\pi} \cdot \vec{\tau}) - g_\omega \not{\omega} - g_\rho \not{\rho} \cdot \vec{\tau}) \psi_2 \\ & + m_0 (\bar{\psi}_2 \gamma_5 \psi_1 - \bar{\psi}_1 \gamma_5 \psi_2)\end{aligned}$$

In a chirally symmetric phase, masses degenerate, $= m_0$.

When chiral symmetry is broken, $\langle \sigma \rangle = \sigma_0 \neq 0$, the masses of the N and N* are split

$$m_\pm = \pm(g_1 - g_2)\sigma_0 + \sqrt{(g_1 + g_2)^2 \sigma_0^2 + m_0^2}$$

“Know” g_1 ; *not* g_2 or m_0 . In vacuum, fit masses to N(939) & N*(1535)

Parity Doubled Urca

After a few days, temperatures for neutron stars are *really* low, ~ 1 MeV.

Usual Urca: $n \rightarrow p + e + \bar{\nu}_e$ Can be hard to do, $k_p^F + k_e^F \geq k_n^F$, k^F = Fermi mom.

Urca emissivity: $Q^{\text{Urca}} \sim G_F^2 (1 + 3 g_A^2) T^6$

But often Urca can't go.

Modified Urca, $n + N \rightarrow p + N + e + \bar{\nu}_e$ $Q^{\text{mUrca}} \sim G_F^2 g_A^2 T^8$

Parity doubled nucleons: PD'd Urca, $n^* \rightarrow p + e + \bar{\nu}_e$

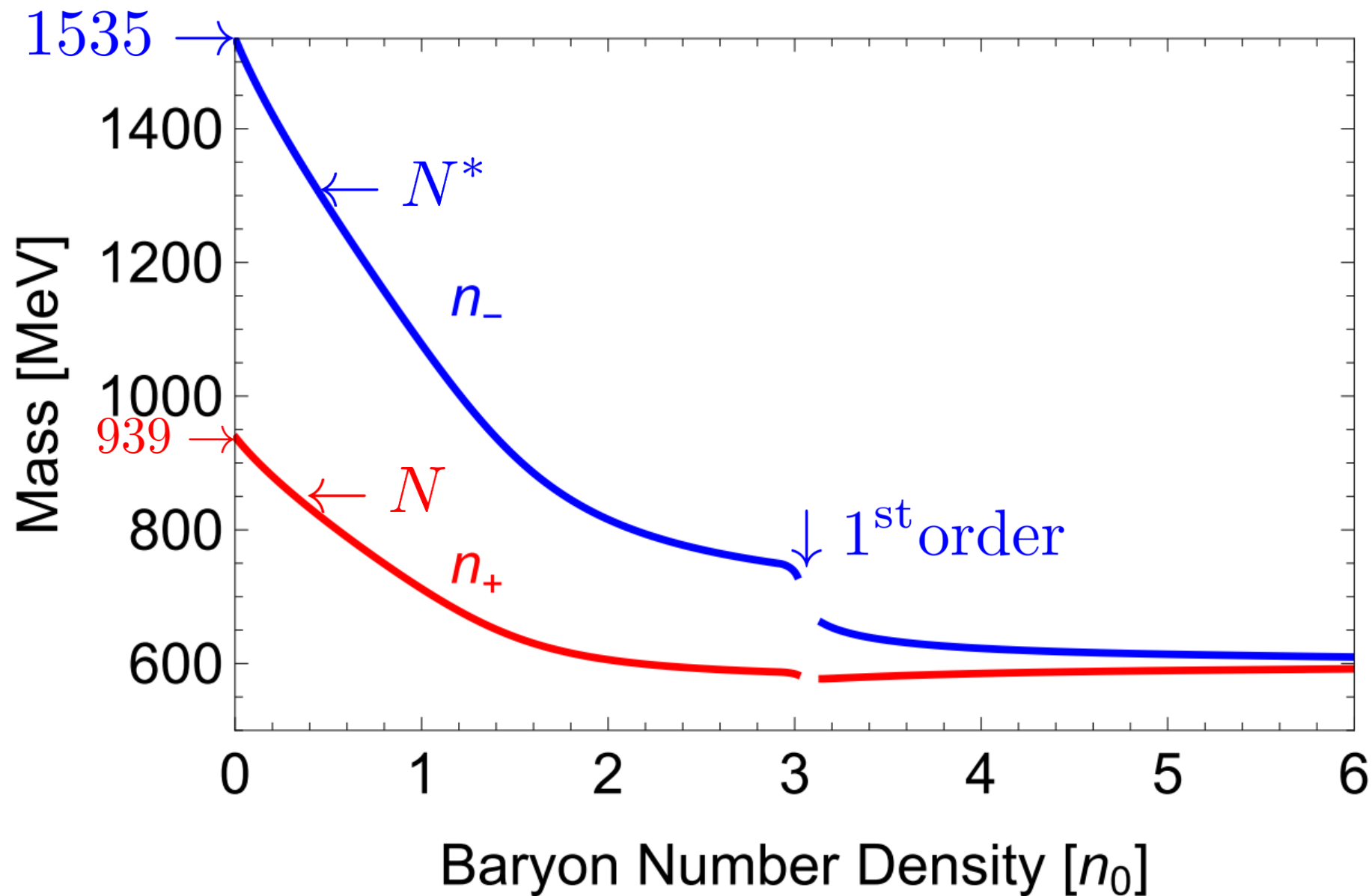
opens up in the *broken* phase! $Q^{\text{PD'd Urca}} \sim G_F^2 (1 + 3 \tilde{g}_A^2) T^6$

Find only in a *narrow* region of density for *heavy* neutron stars.

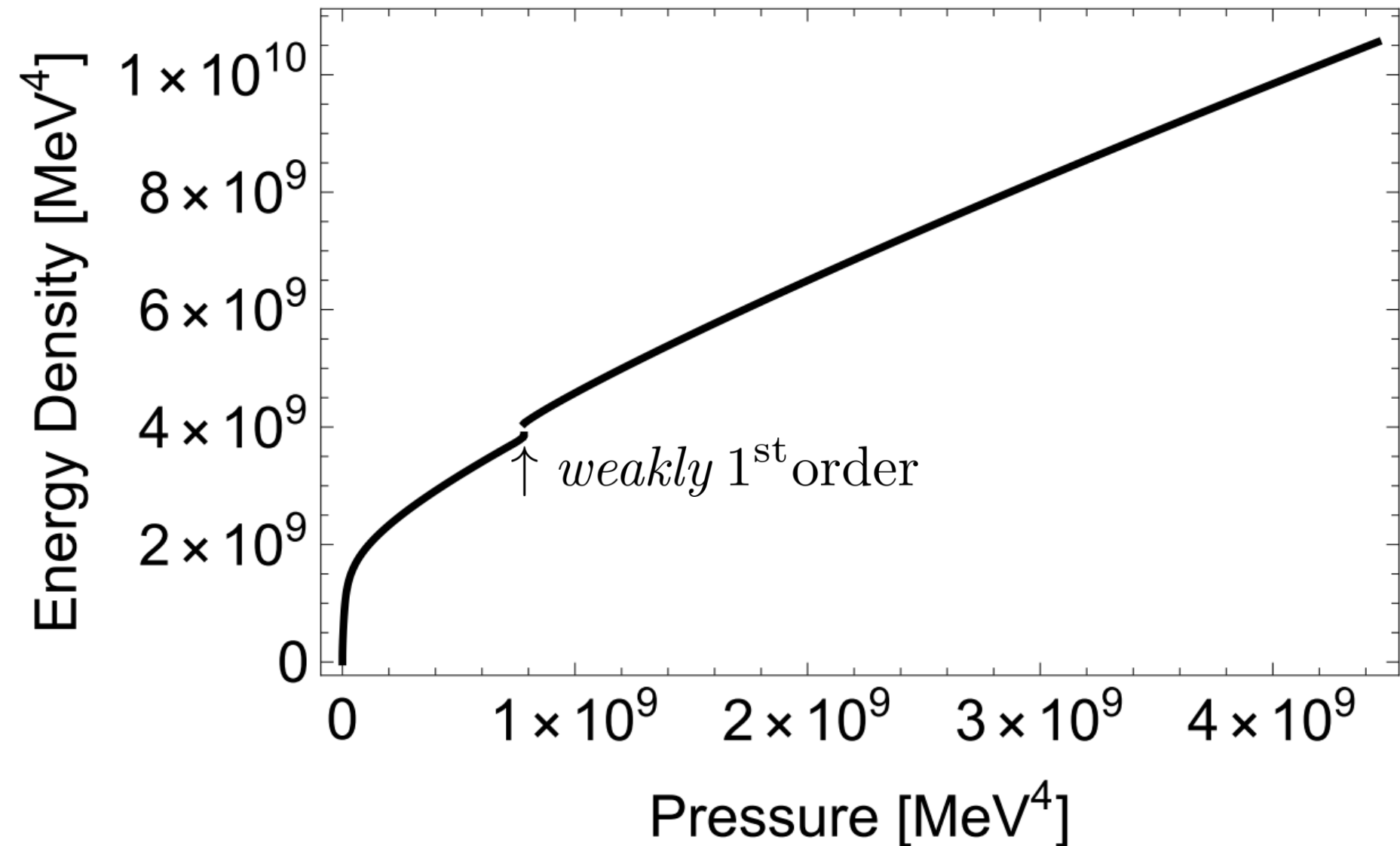
Sensitive to the value of g_A for PD'd nucleons $\tilde{g}_A$;

Kummer, Leupold & von Smekal, [2512.03894](#)

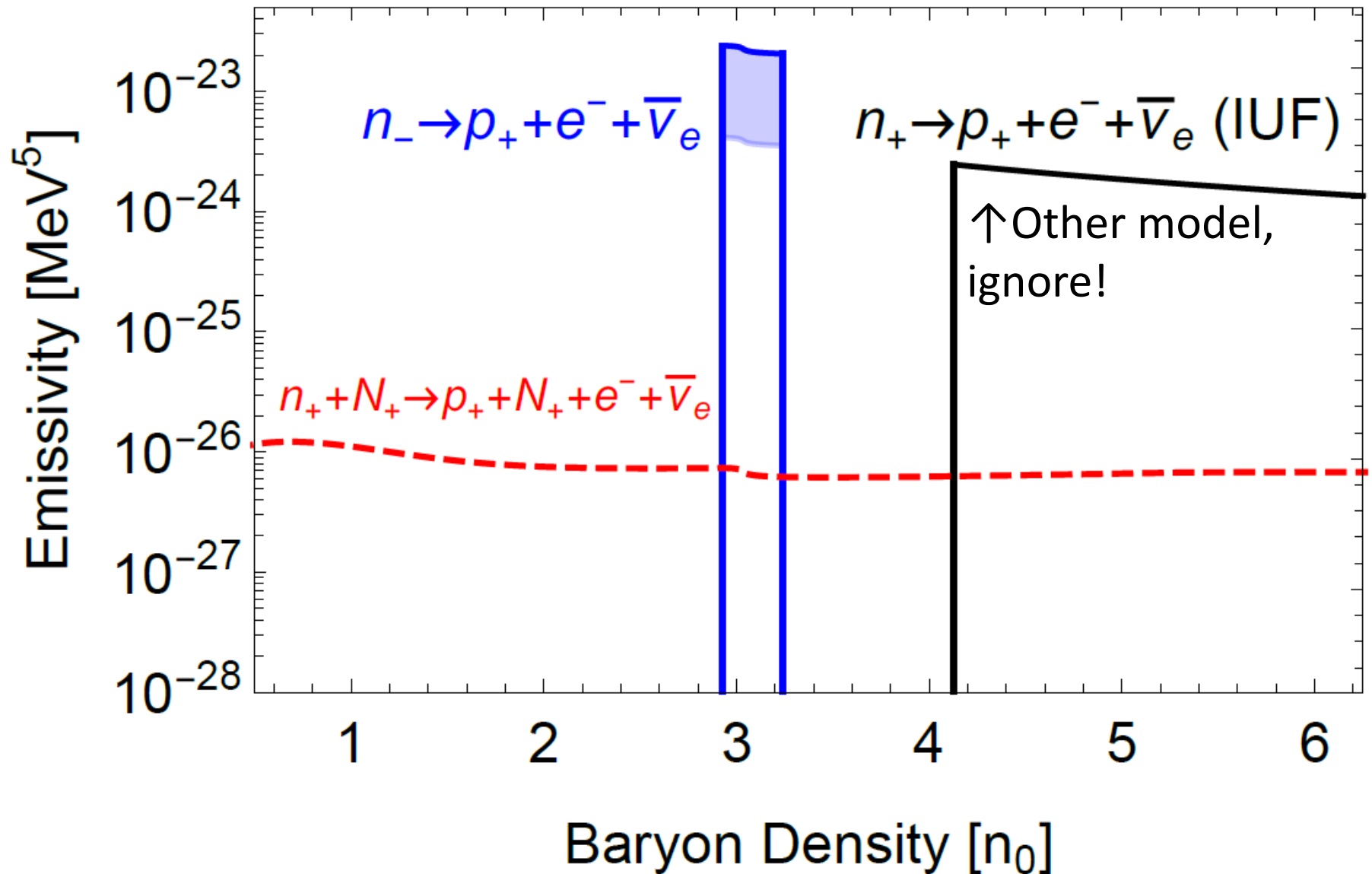
N^* and N masses vs density



Energy vs pressure: *weakly* 1st order



Emissivity for PD'd Urca vs mUrca: “Goldilocks regime”



't Hooft anomalies and the chiral phase transition

Shi Chen, Aleksey Cherman, & RDP, [2603.09977](#)

For three massless flavors at nonzero temperature

The linear sigma model is *wrong*

Predicts 1st, instead of 2nd order chiral transition.

Does not satisfy “‘t Hooft anomalies”.

Applies to the chiral transition at nonzero temperature

The chiral phase transition in QCD is “beyond Landau-Ginzburg”

At least for three flavors, in the chiral limit...



Brookhaven
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Ferromagnets & spin waves

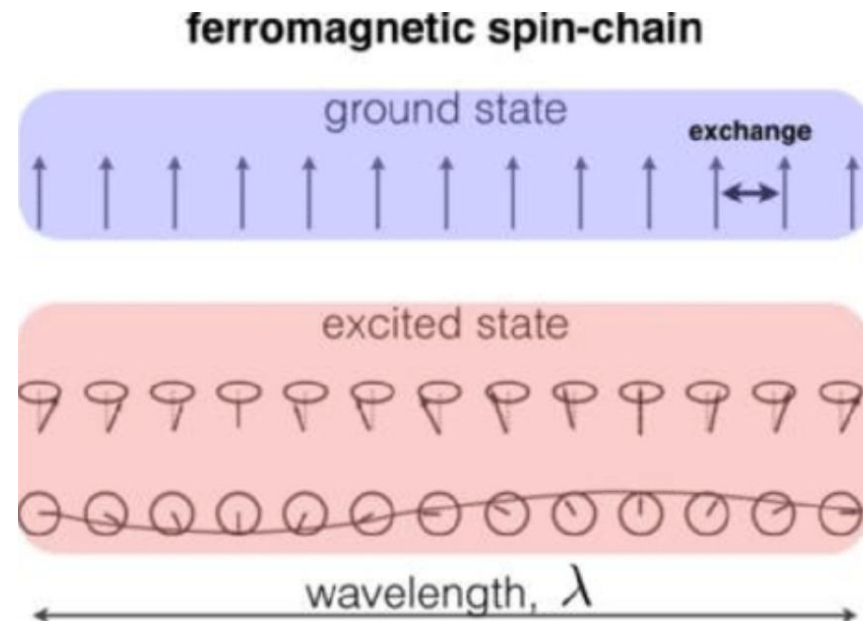
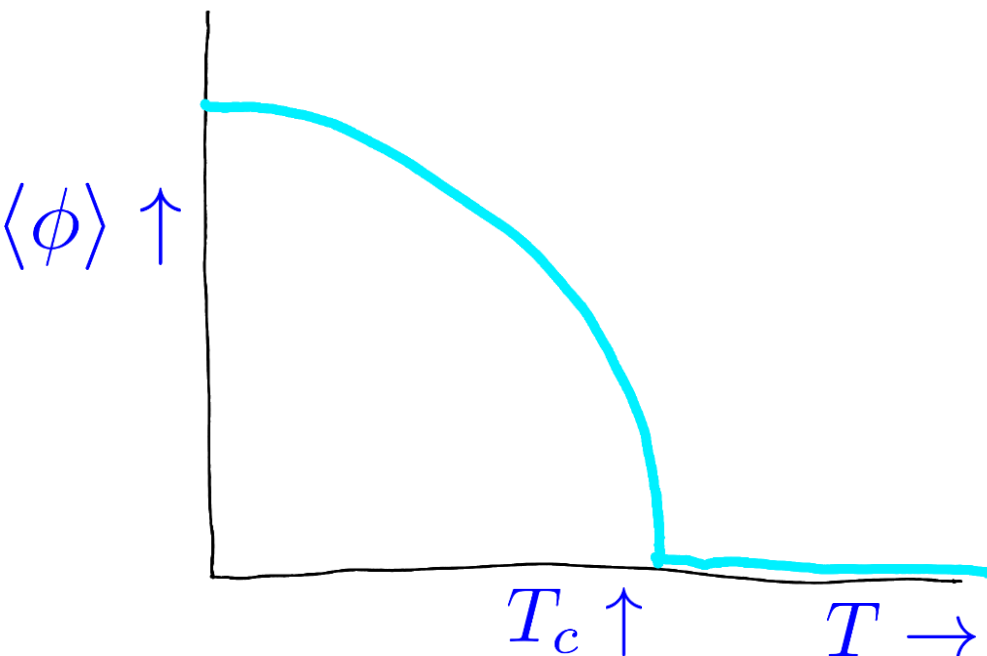
Ferromagnet: spins which like to line up with one another.

Low temperature: magnetization. High temperature: thermal fluc.'s disorder

So the magnetization goes $\downarrow$ as T goes $\uparrow$. Below for 2nd order trans. @ T_c .

If symmetry exact, *massless* excitations: Nambu- Goldstone bosons.

If symmetry almost exact, *light* excitations. Pions light because of approx. chiral sym.



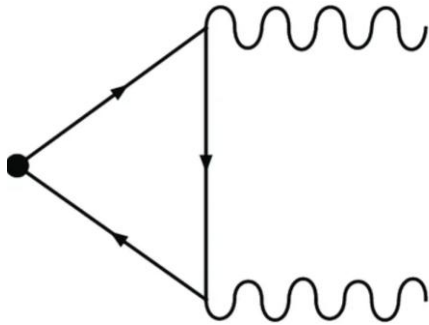
Chiral symmetry in QCD

For N_f flavors of massless quarks, the QCD Lagrangian is invariant under

$$q_L \rightarrow e^{i\theta_V + i\theta_A} U_L q_L ; q_R \rightarrow e^{i\theta_V - i\theta_A} U_R q_R ; U_{L,R}^\dagger U_{L,R} = \mathbf{1}_{N_f}$$

Global chiral symmetry, $G_{cl} = U(1)_V \times U(1)_A \times SU(N_f)_L \times SU(N_f)_R$.

Valid classically. $U(1)_A$ violated by *axial anomaly*, $U(1)_A \rightarrow Z(N_f)$:



$$\partial_\mu J_\mu^5 \sim \text{tr } G_{\mu\nu} \tilde{G}^{\mu\nu} ; J_\mu^5 = \bar{q} \gamma_\mu \gamma_5 q$$

$$Q_{\text{top}} \sim \int d^4x \text{tr } G_{\mu\nu} \tilde{G}^{\mu\nu} = \text{integer}$$

The isosinglet channel is special, and is *not* a symmetry of the quantum theory.

For three flavors, octet of π 's, K's, & η ; singlet is η' .

When chiral symmetry is spontaneously broken, *light* octet, *heavy* singlet.

The linear sigma model

Natural order parameter for chiral symmetry:

$$\Phi = \bar{q}_L q_R \rightarrow e^{i\theta_A} U_L^\dagger \Phi U_R$$

In vacuum, $\langle \Phi \rangle \neq 0$, symmetry then $U(1)_B \times SU(N_f)_V$

$$q_{L,R} \rightarrow e^{i\theta_V} U_V q_{L,R}$$

Construct effective Lagrangian from all terms invariant under the *unbroken* symmetry:

$$\mathcal{L} = \text{tr} |\partial_\mu \Phi|^2 + m^2 \text{tr} \Phi^\dagger \Phi + \lambda_1 (\text{tr} \Phi^\dagger \Phi)^2 + \lambda_2 \text{tr} (\Phi^\dagger \Phi)^2$$

Can't be everything! Terms $\sim \Phi^\dagger \Phi$ are *invariant* under $U(1)_A$!

Anomalous terms

Because of the axial anomaly, *must* include terms $\sim \det \Phi$.

$$\text{Under } G_{\text{cl}} \text{ rotation, } \det \Phi \rightarrow e^{i N_f \theta_A} \det \Phi$$

If $\theta_A = 2 \pi / N_f$, $\det \Phi$ invariant. Origin of $Z(N_f)$ symmetry.

Terms with $\det \Phi$ *must* be included in the effective Lagrangian:

$$\mathcal{L}_A = c_1 \det \Phi + c_2 (\text{tr } \Phi^\dagger \Phi) \det \Phi + c_3 (\det \Phi)^2 + \text{c.c.}$$

These terms are related to “*instantons*”,
which carry topological charge,

$$Q_{\text{top}} \sim \text{tr} G_{\mu\nu} \tilde{G}_{\mu\nu} \neq 0$$

The coefficients of $\det \Phi$ and $\text{tr } \Phi^\dagger \Phi \det \Phi$ are due to $Q_{\text{top}} = 1$;

That $\sim (\det \Phi)^2$ to $Q_{\text{top}} = 2$. RDP & Rennecke, [1910.14052](#)

In vacuum, c_1 (or c_2 or c_3) *must be* large to make the η ' heavy.

...and the chiral phase transition

The anomaly term depends *crucially* upon the number of flavors:

$$\det \Phi \sim \Phi^{N_f}$$

Two flavors: a mass term,
splits the η from the σ , and the a_0 's from the π 's

$$m_\eta^2 \sim m_\sigma^2 + c_1$$

Three flavors: again, splits the η' from the σ

$$m_{\eta'}^2 \sim m_\sigma^2 + c_1 \langle \Phi \rangle$$

Chiral phase transition: if second order, $\langle \Phi \rangle \rightarrow 0$

Two flavors: $\sigma + \pi = O(4)$ go critical, a_0 's and η stay massive.

Three flavors: $\det \Phi$ is cubic. *Must be first order if $c_1 \neq 0$ Unavoidable.*

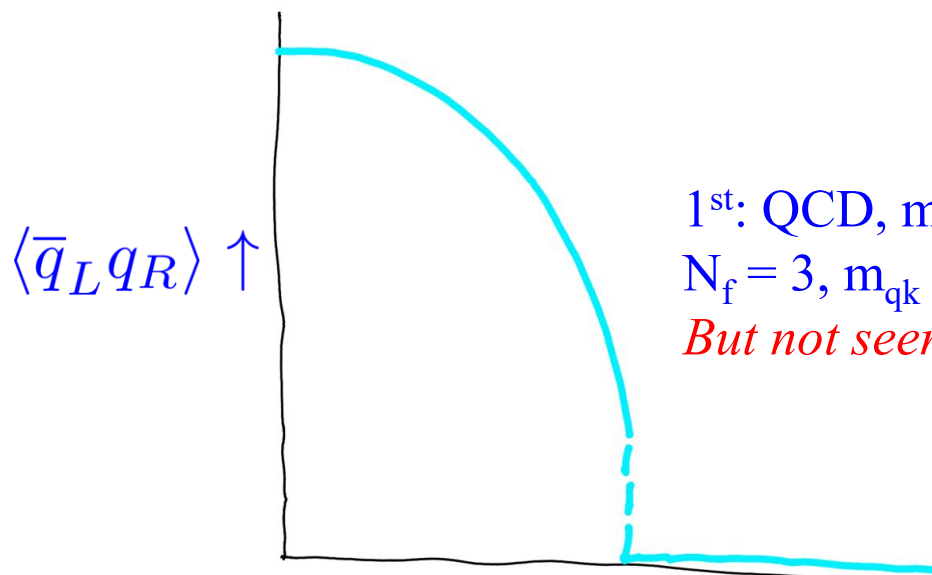
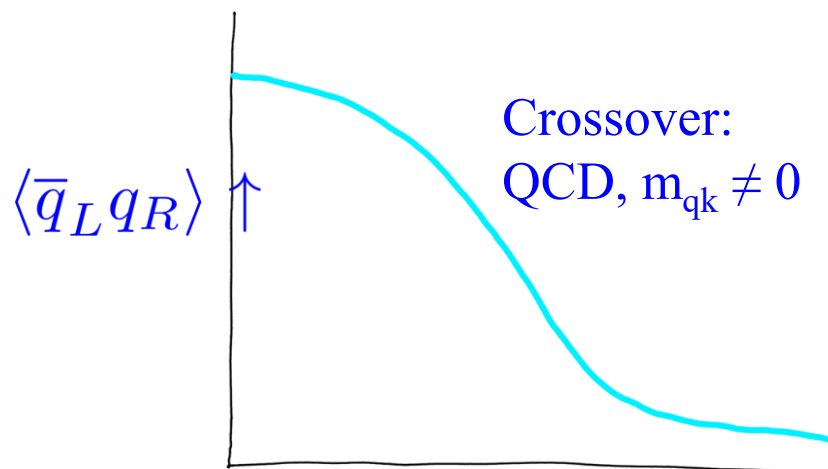
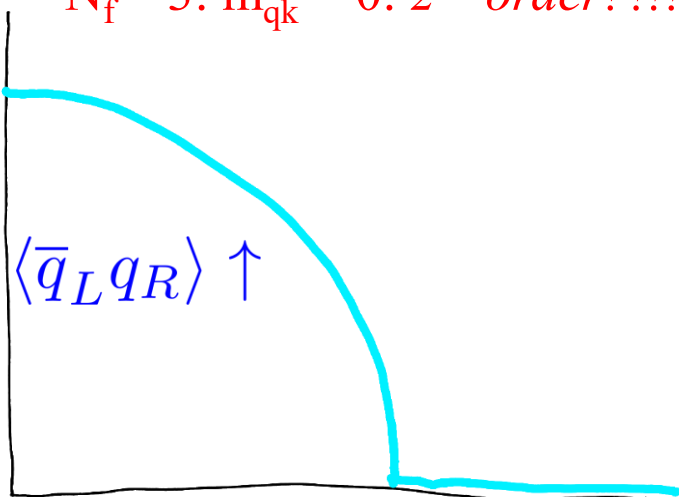
RDP & Wilczek, '84.

Phase transitions for the chiral transition: 2^{nd} order for 3 flavors?

$N_f = 2: m_{qk} = 0, 2^{nd}$ order

$N_f = 3: m_{qk} = 0: 2^{nd}$ order?!!!!

x — axis : T or μ

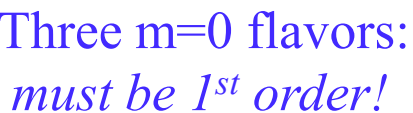


X

2 massless flavors: 2nd order. $QCD = crossover$

Two $m=0$ fl.'s, $\rightarrow$
 2^{nd} order (lattice)

← pure gauge,
1st order .
due to *cubic* invariant
from Z(3) global sym.


$$m_u = m_d \rightarrow$$

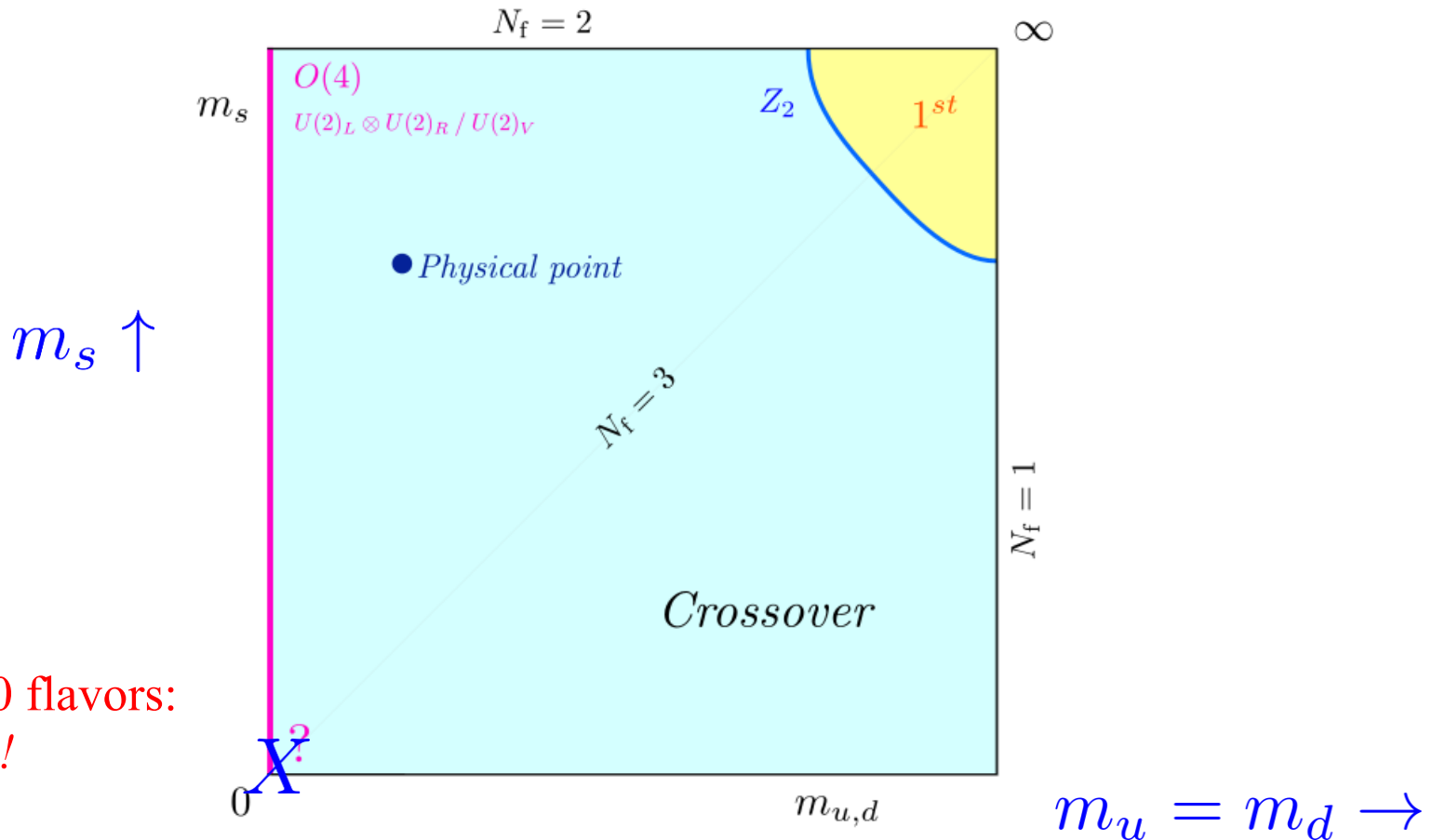
Lattice: *no* sign of 1st order for 3 massless flavors!

“Goethe” phase diagram. Cuteri, Philipsen, Sciarra, [2112.11107](#):

Chiral transition appears to be 2nd order for 3 $m=0$ flavors! *As if $c_1 = 0$ at T_χ !*

Restore $U(1)_A$ at T_χ ? RDP & Rennecke, [2401.06130](#)

Phase diagrams with c_1, c_2, c_3 : Giacosa, Kovacs², RDP & Rennecke, [2410.08185](#)



't Hooft anomalies

First, *completely* ignore $U(1)_A$, always broken (vanishes as $T \rightarrow \infty$: $c_1 \sim 1/T^7$, etc)

Global symmetry group is

$$\mathcal{G}_{\text{qu}} = \frac{U(1)_V \times SU(N_f)_L \times SU(N_f)_R}{Z(3) \times Z(N_f)}$$

After chiral symmetry breaking,

$$\mathcal{H} = \frac{U(1)_V \times SU(N_f)_V}{Z(N_f)}$$

Boundary conditions in imaginary time:

in pure gauge theory, gauge transf.'s

Ω can have $Z(3)$ “twist”:

$$\Omega(1/T) = e^{2\pi i/3} \Omega(0)$$

Gluons are invariant: $A_\mu(1/T) \rightarrow e^{-2\pi i/3} A_\mu(0) e^{+2\pi i/3} = A_\mu(0)$

~ Polyakov loop, order parameter for deconfinement. “1-form” order parameter.

Quarks (in fund. rep.) are *not*:

$$q(1/T) = e^{2\pi i/3} q(0)$$

Roberge-Weiss transition

Have $U(1)_B$, introduce chemical potential μ .

Boundary conditions now $q(1/T) = e^{\mu/T} q(0)$

Thus for $Z(3)$ twisted b.c.,

$$q(1/T) = e^{2\pi i/3 - \mu/T} q(0) = q(0) ; \text{ if } \mu = \frac{2\pi T}{3} i$$

Roberge & Weiss '88. Thus the partition function is invariant under this shift:

$$\mathcal{Z} \left(\mu + \frac{2\pi}{3} i \right) = \mathcal{Z}(\mu)$$

N.B.: this is an *imaginary* shift in the chemical potential!

Who cares about an imaginary chemical potential?



't Hooft anomalies

't Hooft '79....Cordova, Freed, Lam, Seiberg, [1905.09315](#) ; [1905.13361](#)

Yonekura, [1901.08188](#) ; Nishimura & Tanizaki, [1903.04014](#);

Kobayashi, Yokokura & Yonekura, [2305.01217](#)

Add background *gauge* fields for the *global* symmetries: A_L^μ, A_R^μ : *not* dynamical!

t Hooft anomaly: $U(1)_V \times SU(N_f)_{L,R} \times SU(N_f)_{L,R}$. Constant (imag.) μ like A_0 for $U(1)_V$

$$\mathcal{Z}(\mu + 2\pi i/3, A_L, A_R) = \mathcal{Z}(\mu, A_L, A_R) \mathcal{J}(A_L) \mathcal{J}^{-1}(A_R)$$

$$\mathcal{J}(A) = \exp \left(\frac{i}{4\pi} \int d^3x \epsilon^{\alpha\beta\delta} \text{tr} \left(A_\alpha \partial_\beta A_\delta - \frac{2}{3} i A_\alpha A_\beta A_\delta \right) \right)$$

↑ Chern-Simons term for the background gauge fields.

Computed perturbatively in the UV. **Must hold for *any* effective theory in the IR;**

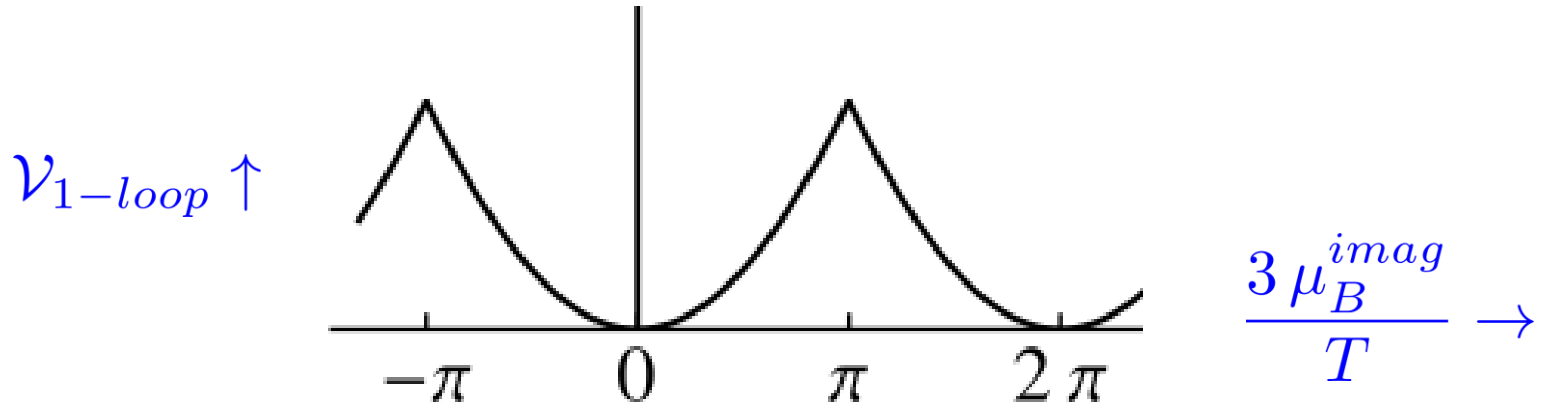
't Hooft: introduce “spectator fermions”

***Strong* constraint on effective theories!** “Anomaly in space of coupling constants”

High T: 1st order transition

At high T, restore chiral symmetry, can compute in perturbation theory.

Imaginary chemical potential = constant, real A_0 field. *Easy* to compute @ 1-loop order!



Roberge & Weiss '88: 1st order transition at $3 \mu_B^{imag} = \pi T$

Quark density imaginary, flips sign.

One way to satisfy 't Hooft anomaly is by 1st order transition.

$\mu_B^{imag} = 0$ *not* analytically connected to $\mu_B^{imag} = 2 \pi T/3$

RW transition exists at high T, presumably ends at some nonzero T. When?

't Hooft anomaly at *low* T

At low T, chiral symmetry is broken, $\Phi = \Omega \phi_0$; $\Omega^\dagger \Omega = \mathbf{1}$

Can rewrite the effective Lagrangian with the Goldstone bosons, the Ω 's

$$\mathcal{S}_{\text{GB}} = \int d^3x \left(\frac{1}{g^2} \text{tr} |\partial_\mu \Omega|^2 + \mu \frac{\epsilon^{ijk}}{24\pi^2} \text{tr}(B_i B_j B_k) \right) ; B_i = \Omega^\dagger \partial_i \Omega$$

The second term is baryon number, related to $\pi_3(\text{SU}(N_f)_V)$.

If μ_B shifts by $2\pi i/3$, $\exp(-S_{\text{GB}})$ shifts by $\exp(2\pi i)=1$.

Gauging S_{GB} , this generates the correct anomaly (Witten '84)

Anomalies arise from fermion loops, *and* bosons with topological terms

In linear sigma model:

$$Q = \int d^3x \epsilon^{ijk} \text{tr}(\tilde{B}_i \tilde{B}_j \tilde{B}_k) ; \tilde{B}_i = \Phi^\dagger \partial_i \Phi$$

But not topological! Q not baryon number, just operator with high mass dimension

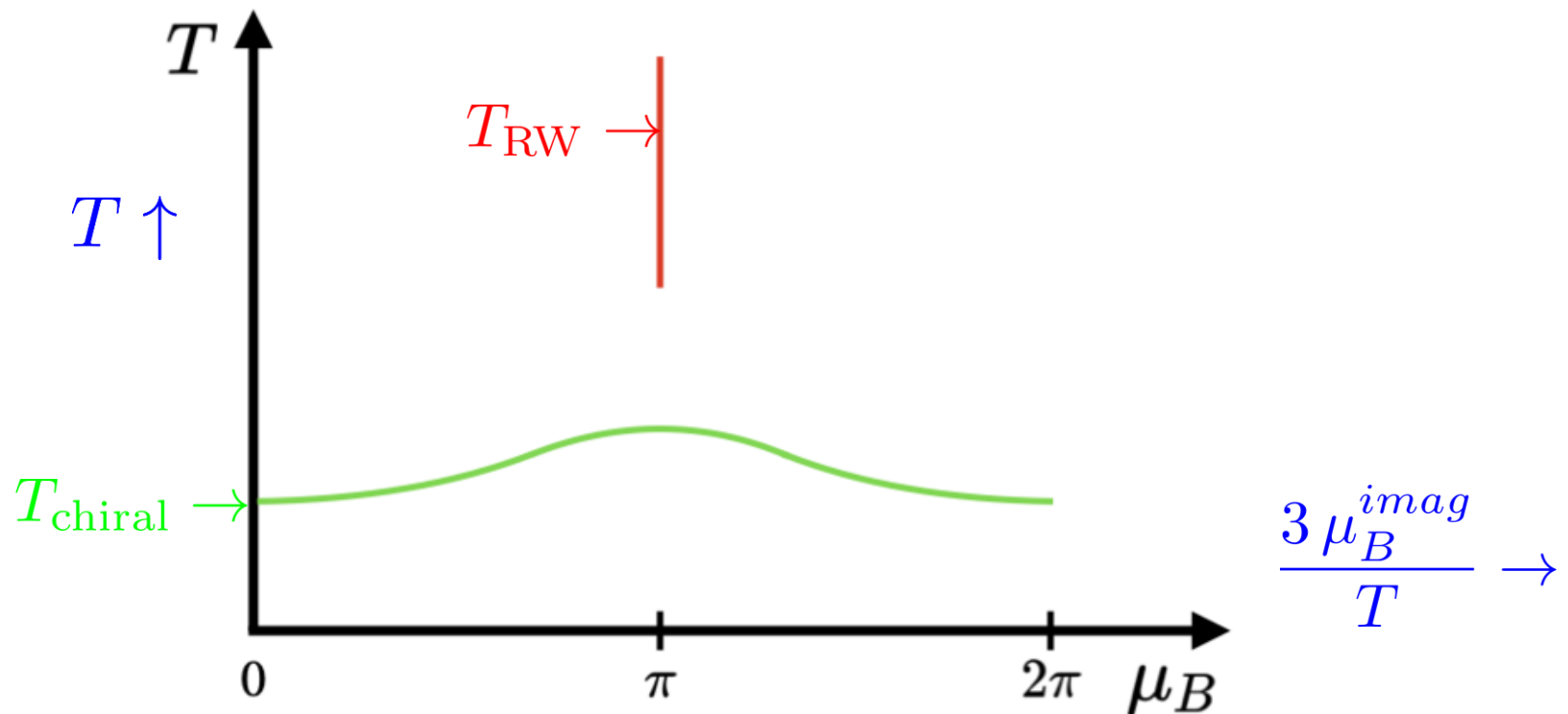
Excluded: $T_{RW} > T_{ch}$.

RW transition: 1st order at $\mu_B^{imag} = \pi T$ for high T , assume same at lower T .

How does the RW line end? *Not* above T_{chiral} !

Linear model does not satisfy 't Hooft anomaly. True for any # flavors.

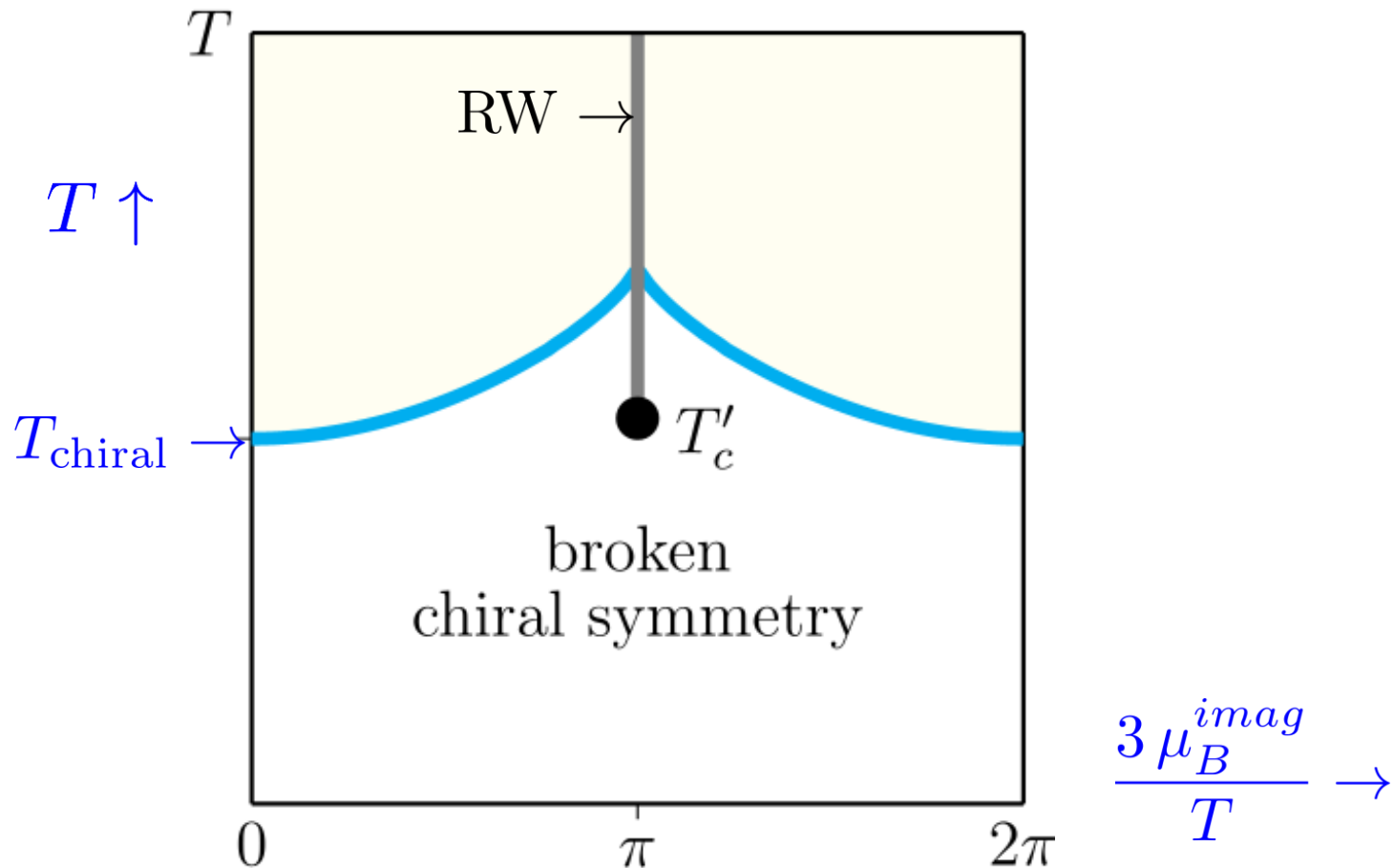
Kobayashi, Yokokura & Yonekura, [2305.01217](#) . Three (other) phase diagrams possible.



Landau: $T_{RW} < T_{ch}$.

The Roberge-Weiss transition can end *below* the chiral, at standard critical endpoint.

Landau works: in symmetric phase, linear model + quarks for RW transition.

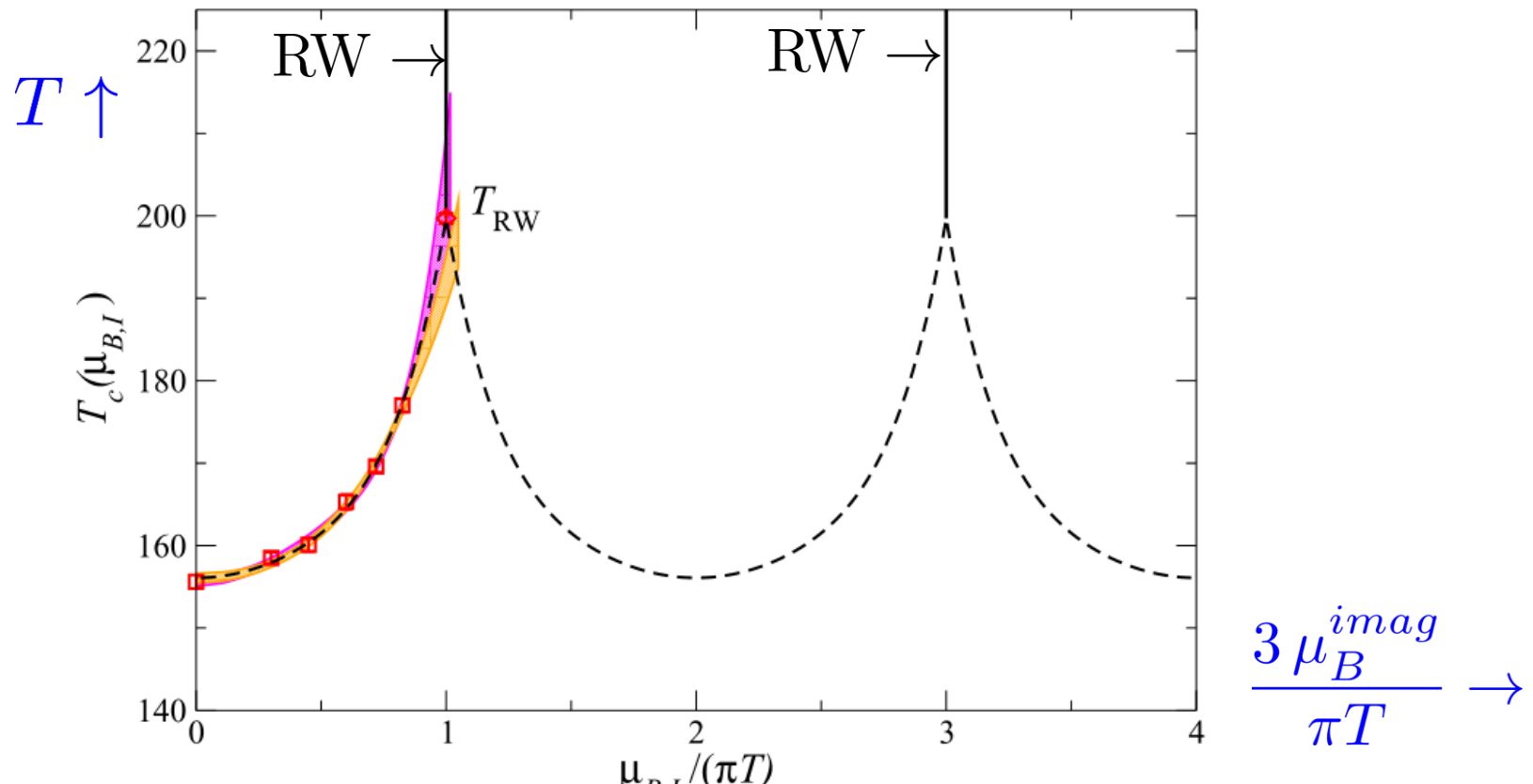


Lattice, 2+1 flavors: $T_{RW} > T_{ch}$.

Easy for lattice to compute at imaginary chemical potential, as real A_0 .

Lattice: Bonati, D'Elia, Mariti, Mesiti, Negro [1602.01426](#), Cuteri+...[2205.12707](#)

QCD, 2+1 flavors: *crossover* at $\mu=0$, remains crossover until $T_{RW} > T_{ch}$.



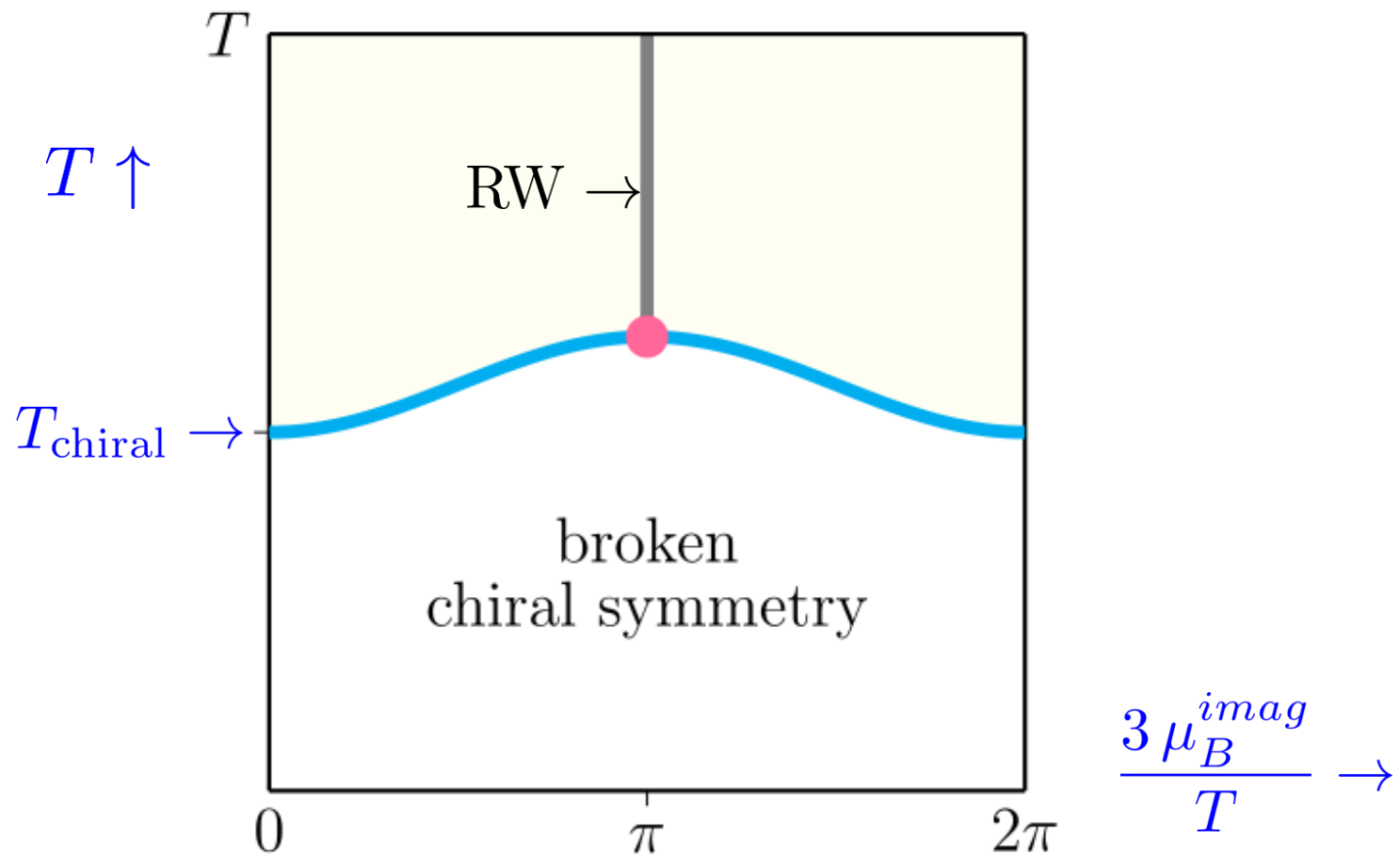
$$\text{DQCP: } T_{\text{RW}} = T_{\text{ch}} .$$

Back to chiral limit. RW transition can end on the critical line for the chiral transition.

1st order RW line ends at Z(2) critical endpoint. *Not* universality class of chiral trans.

“Deconfined quantum critical point”

“Beyond Landau”: Landau *cannot* describe change in universality class.

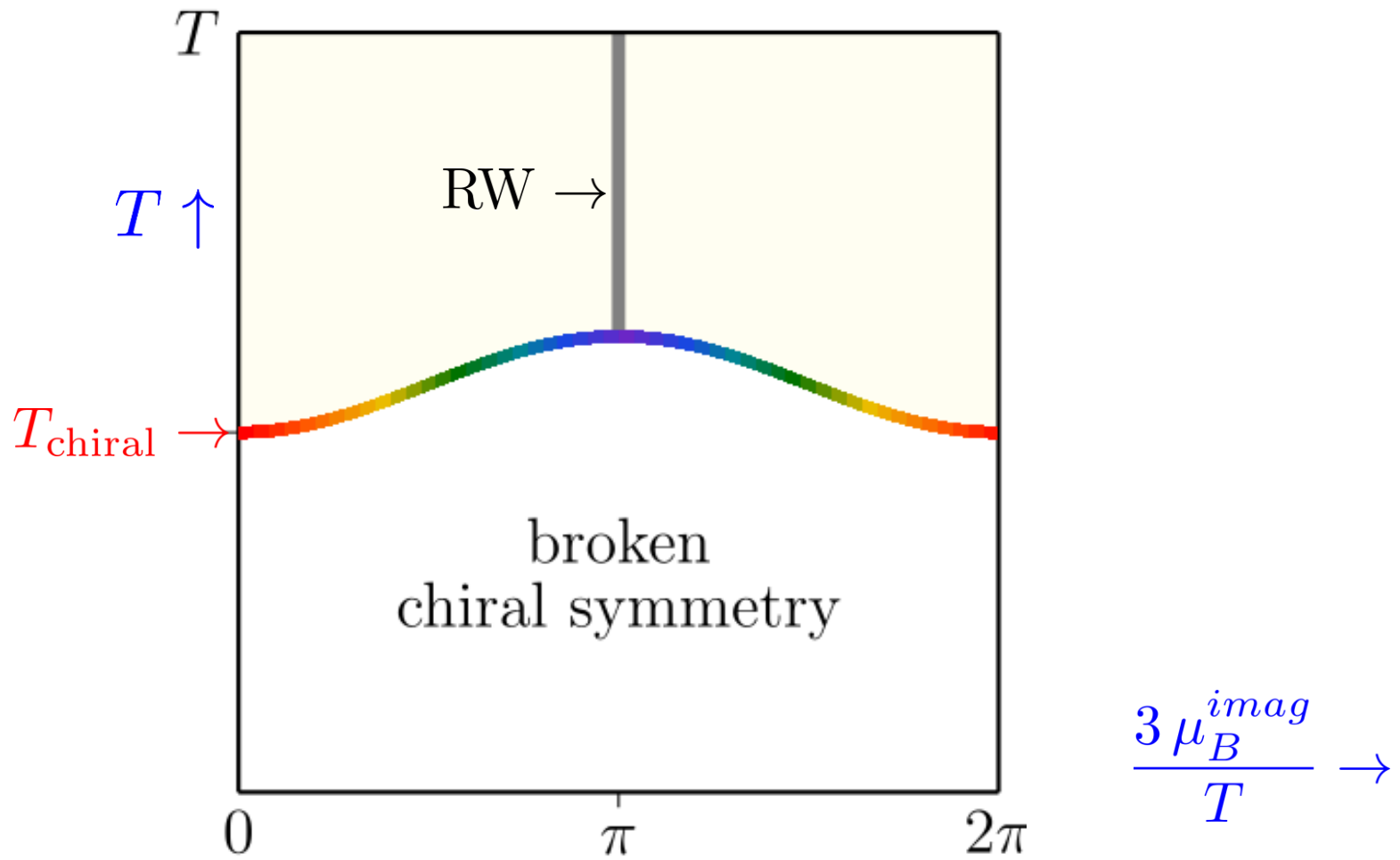


Continuous manifold: $T_{\text{RW}} = T_{\text{ch}}$.

Roberge-Weiss ends on critical line, but universality class changes *continuously* with μ_B

Requires that the baryon density is *exactly* marginal for *all* μ_B .

Beyond Landau. No known example without SUSY. Unusual, possible.



Conclusions

For two flavors, chiral transition is 2nd order, ok with linear model.

For three flavors, in the chiral limit trans. of 2nd order, *not* 1st order, as in linear model.

Linear model fails to describe 't Hooft anomaly; ok if 1st order RW intervenes.

Extension to nonzero chemical potential?

Effective theory for three flavors? What about $\mu_B \neq 0$?

Instead of linear σ model, going *down* from $4 - \varepsilon$ to 3 dimensions,

work with a non-linear σ model, going *up* from $2 + \varepsilon$ dimensions to 3.

Really? A non-linear model?

Certainly must add quarks to any effective model to get the RW transition right.

Quarkyonic models, which always have quarks, appear inescapable.